

The image shows three large, cylindrical scientific instruments, identified as UVES LP, mounted on a ship's deck. The instruments are arranged in a row from left to right. The central instrument is the most prominent, showing its complex structure and various panels. The background features a sunset over the ocean, with the sun low on the horizon, casting a warm glow. The sky is a deep purple and blue. The instruments are dark in color, possibly black or dark grey, with some metallic components visible. The overall scene is dramatic and highlights the scale of the equipment.

***The UVES LP for Testing Fundamental
Physics: Status and Dark Side Constraints***

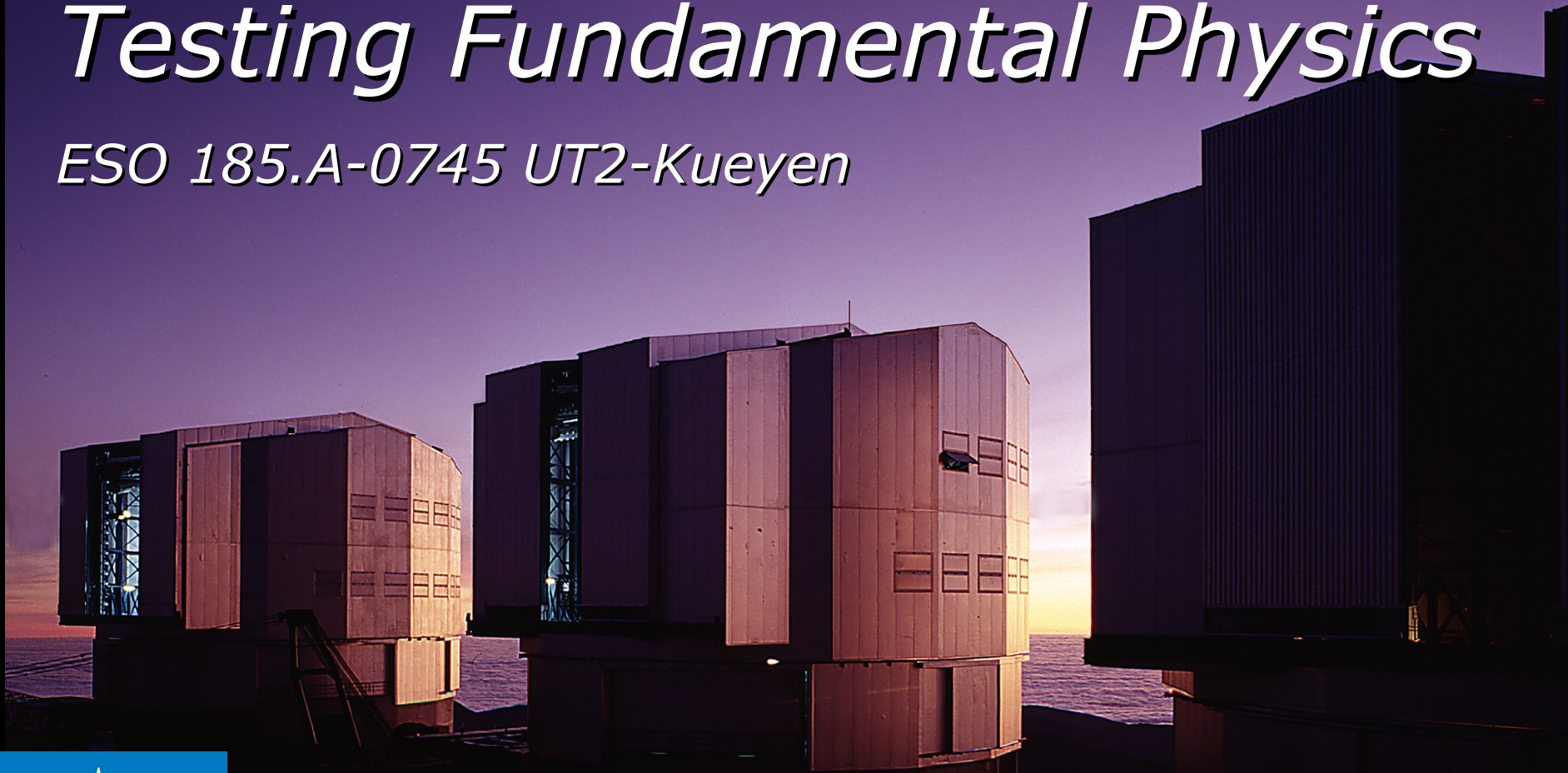
Carlos Martins and the CAUP Dark Side Team

So What's Your Point?

- We all know that fundamental couplings run with energy
- Moreover, in many (or arguably most?) models they will equally naturally roll in time and ramble in space
- Therefore astrophysical (and local) tests of their stability provide us with optimal probes of fundamental cosmology
- In this talk I will present the state-of-the-art astrophysical measurements, and highlight some of their implications

The UVES Large Program for Testing Fundamental Physics

ESO 185.A-0745 UT2-Kueyen



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LP Plan & Goals

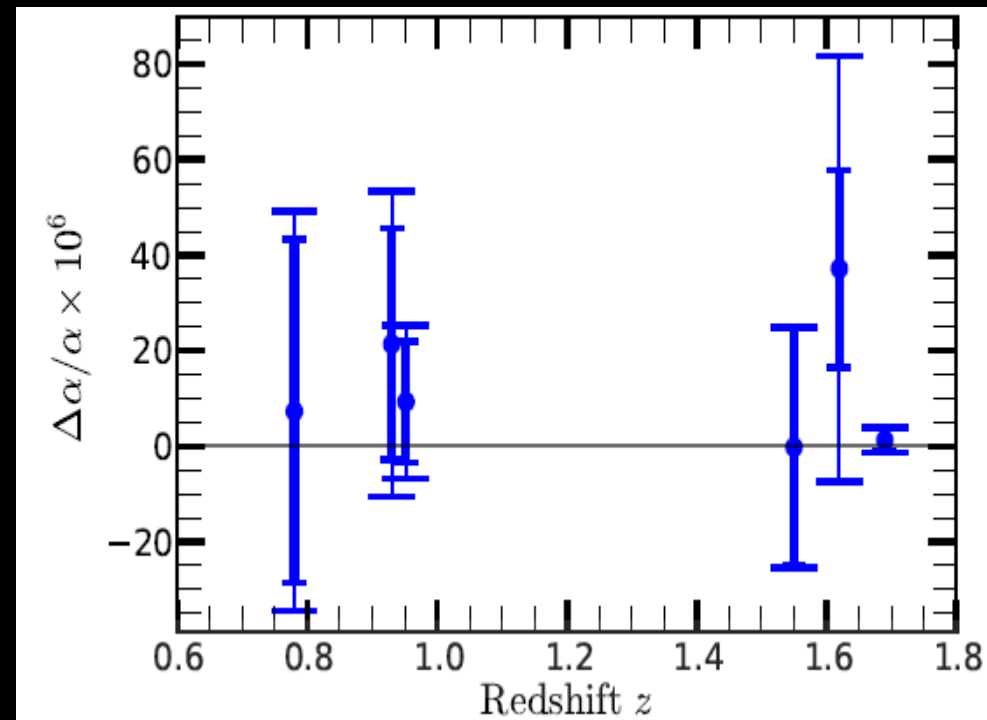
- Only program dedicated to varying couplings, with optimized data acquisition and reduction: ca. 40 nights in 2010-13
 - Calibration lamps attached to science exposures (in same OB): don't reset x-disperser encoding position for each exposure
 - Observe bright (mag 9-11) asteroids at twilight, to monitor radial velocity accuracy of UVES and the optical alignments
- $R \sim 60000$, $S/N \sim 100$; potential accuracy is 1-2ppm/system, where photon noise and calibration errors are comparable
 - Our goal: 2ppm per system, 0.5ppm for full sample
 - All active groups involved, multiple blind independent analyses
- Target selection discussed in [*Bonifacio et al. 2014*]
 - 13 targets for α , 2 targets for $\mu = \text{mp/me}$
 - Already out: HE2217-2818 HE0027-1836, HS1519+1919
 - Raw data in ESO archive, reduced data to come – have fun!

Understanding the Data

- HE2217-2818, $z_{\text{abs}} \sim 1.69$:

$$\Delta\alpha/\alpha = 1.3 \pm 2.4_{\text{sta}} \pm 1.0_{\text{sys}} \text{ ppm}$$

- Paper I: P. Molaro et al., *A&A* 555 (2013) A68
- Dipole fit: $(3.2-5.4) \pm 1.7$ ppm depending on model; our measurement does not confirm this, but is not inconsistent with it either



- HE0027-1836, $z_{\text{abs}} \sim 2.40$: $\Delta\mu/\mu = -7.6 \pm 8.1_{\text{sta}} \pm 6.3_{\text{sys}} \text{ ppm}$

- Paper II: H. Rahmani et al., *MNRAS* 435 (2013) 861
- Identified wavelength-dependent velocity drift (corrected with bright asteroid data)

- Bottleneck: intra-order distortions (~ 200 m/s) & long-range distortions on UVES, discussion in Paper IV [Whitmore et al.]

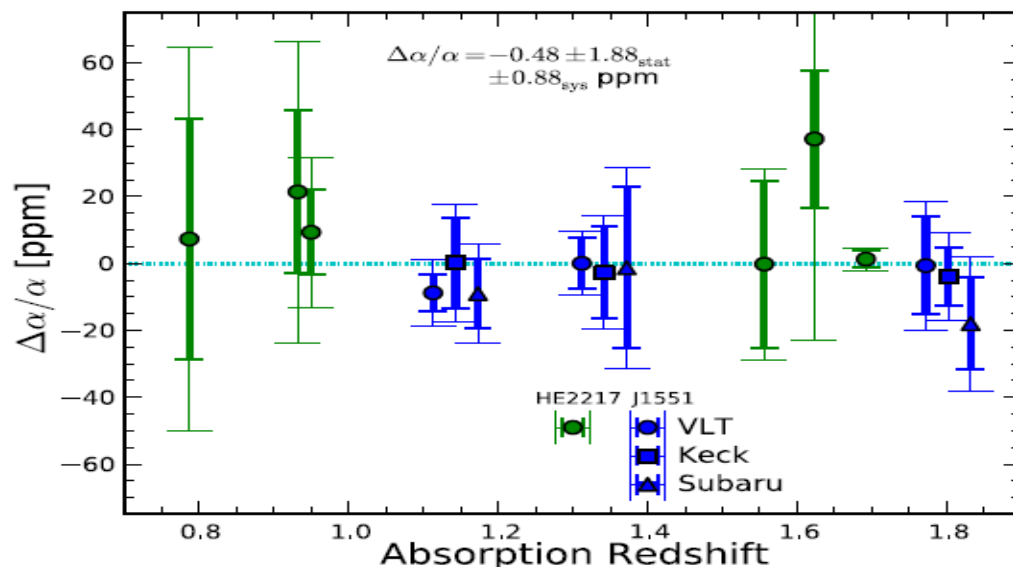
- Also identified in HARPS and Keck-HIRES

A Triple Check of Distortions

- HS1519+1919: 3 absorbers at $z_{\text{abs}} \sim 1.1, 1.3 \text{ \& } 1.8$, observed with 3 top optical telescopes: $\Delta\alpha/\alpha = -5.4 \pm 3.3_{\text{stat}} \pm 1.5_{\text{sys}}$ ppm
 - Paper III: T. Evans et al., MNRAS 445 (2014) 128
 - Directly comparing spectra and 'super-calibrating' with asteroid and iodine-cell tests, allows removal of long-range distortions

Absorption Redshift	Keck/HIRES				$\Delta\alpha/\alpha \pm \sigma_{\text{stat}} \pm \sigma_{\text{sys}}$ [ppm]				Subaru/HDS				Absorber Average		
	χ^2_{ν}	χ^2_{ν}	χ^2_{ν}	χ^2_{ν}	VLT/UVES	χ^2_{ν}	VLT/UVES	χ^2_{ν}	χ^2_{ν}	χ^2_{ν}	χ^2_{ν}	χ^2_{ν}	χ^2_{ν}	χ^2_{ν}	
$z_{\text{abs}} = 1.143$	+0.20	13.63	3.97	1.18	-8.80	5.60	4.36	1.45	-9.04	10.41	4.34	1.59	-7.49	4.63	3.02
$z_{\text{abs}} = 1.342$	-2.77	13.71	3.16	1.20	+0.02	7.64	1.85	1.53	-1.29	24.04	6.04	1.23	-0.70	6.43	1.55
$z_{\text{abs}} = 1.802$	-3.92	8.61	4.69	0.75	-0.66	14.65	4.54	0.98	-17.98	13.67	6.45	0.76	-6.42	6.52	3.16
Weighted mean	-2.64	6.43	2.54	-	-4.71	4.31	2.36	-	-11.20	7.83	2.44	-	-5.40	3.25	1.53

- Current status: compatible with null result and dipole...
 - Full sample analysis ongoing
 - Papers IV-VII should be appearing soon



Dark Energy & Varying Couplings

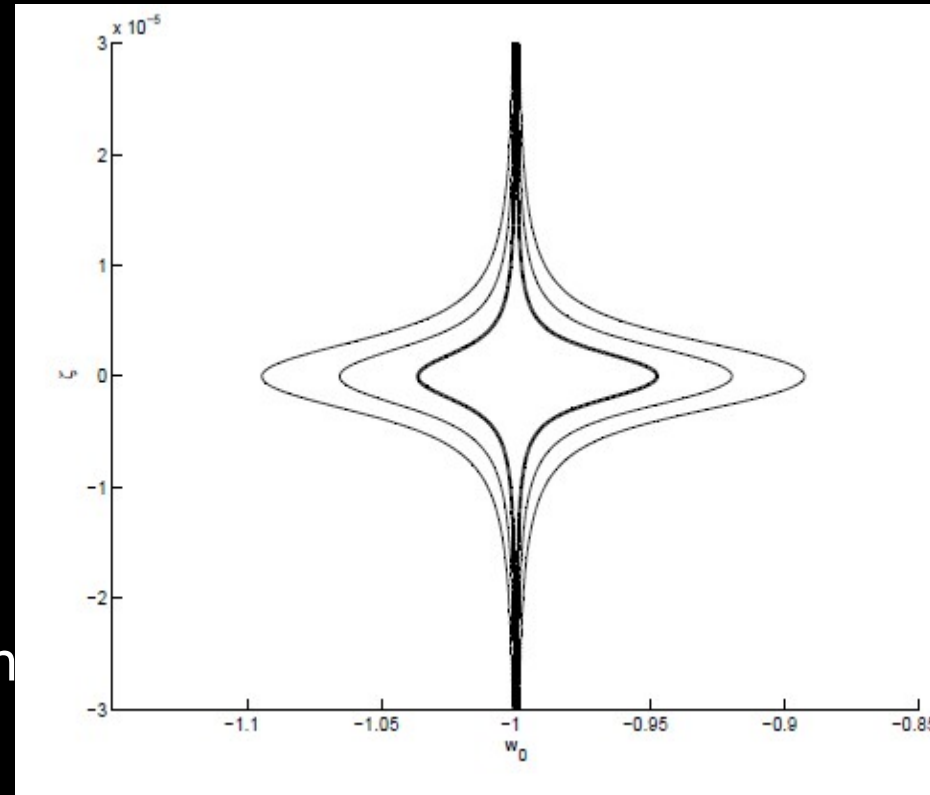
- Universe dominated by component whose gravitational behavior is similar to that of a cosmological constant
 - A dynamical scalar field is (arguably) more likely
- Such a field must be slow-rolling (mandatory for $p < 0$) and be dominating the dynamics around the present day
- Couplings of this field will lead to potentially observable long-range forces and varying 'constants' [Carroll 1998]
 - These measurements (whether they are detections of null results) will constrain fundamental physics and cosmology
 - This ensures a 'minimum guaranteed science'

Taxonomy: Class I

- If the same degree of freedom is responsible for dark energy and varying α , its evolution is parametrically determined

- Current QSO + Clocks + Cosmo marginalized constraints are *[Martins & Pinho 2015]*

- $|\zeta| < 5 \times 10^{-6}$ (2 sigma)
- $|1+w_0| < 0.06$ (3 sigma)
- Atomic clocks currently provide the tightest constraint...
- ...but this will likely change when further LP results come out



- ALMA, ESPRESSO and ELT-HIRES will map the dark side out to $z \sim 4$ *[Amendola et al. 2012, Leite et al. 2014]*
 - Key synergies with other probes (cf. Ana Catarina Leite's talk)
 - For a roadmap in an E-ELT context see *[Martins 2014]*

Euclid & Varying α



- The weak lensing shear power spectrum + Type Ia SNe can constrain Class I models

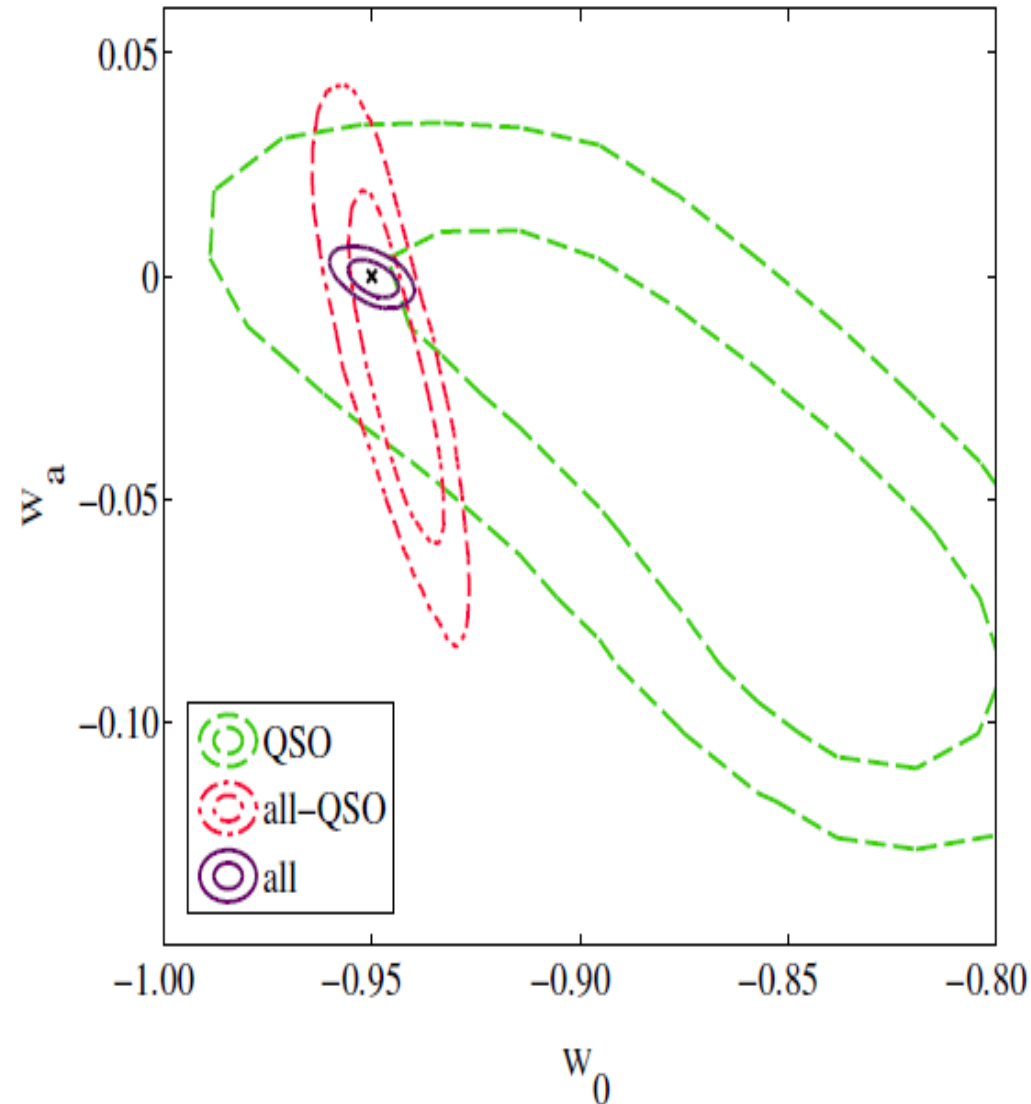
- ...with external datasets

- Example for a CPL fiducial

- Euclid WL + DESIRE SN Ia data [Astier et al. 2014]
- ELT spectroscopic data (+ atomic clock prior)

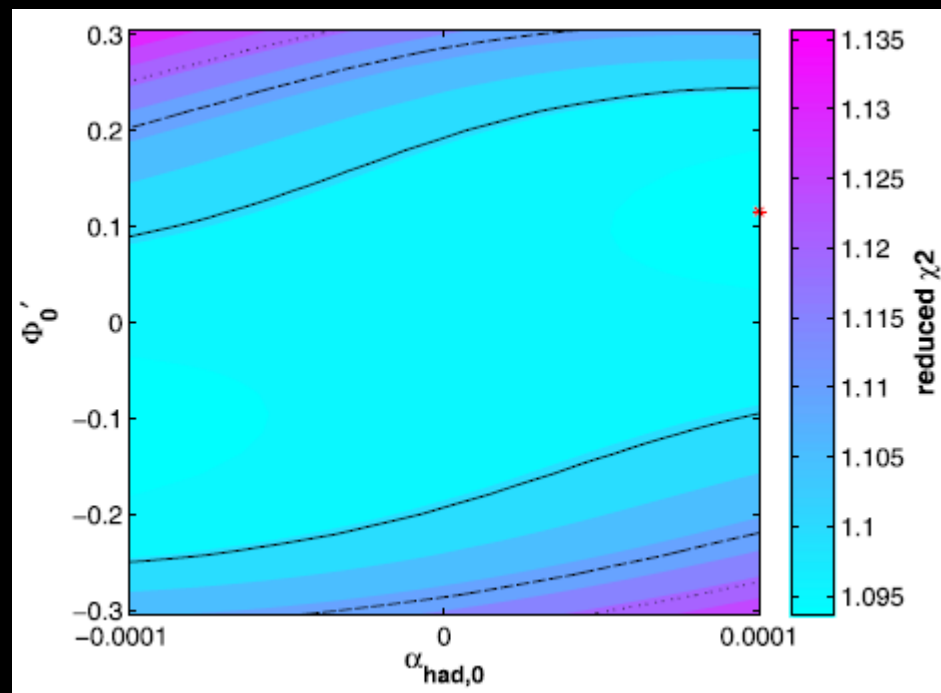
- For a full analysis see [Calabrese et al. 2014]

- Key synergies between Euclid and the various E-ELT instruments are currently being quantified



Taxonomy: Class II

- Models where α field does not provide all dark energy can be identified via $w(z)$ consistency tests [Vielzeuf & Martins 2012]
 - BSBM models [Sandvik et al. 2002, Leal et al. 2014]
 - Runaway dilatons [Damour et al. 2002, Martins et al. 2015]
- Even if this degree of freedom does not dominate the universe at low z , it can bias cosmological parameter estimations
 - Several effects already quantified within Euclid Consortium [Calabrese et al. 2014, Avgoustidis et al. 2014]
 - Tests of the CMB temperature-redshift relation will be a key external dataset [Avgoustidis et al. 2012, de Martino et al. 2015, Luzzi et al. 2015]



So What's Your Point?

- Observational evidence for the acceleration of the universe demonstrates that canonical theories of cosmology and particle physics are incomplete, if not incorrect
 - Fundamental coupling stability is optimal probe of new physics
- The story so far: nothing is varying at $\sim 10^{-5}$ level, already a very significant constraint (stronger than the Cassini bound)
 - At 10^{-6} level things are currently less clear...
 - ...but significant improvements are coming
- Forthcoming instruments will lead to a new generation of precision consistency tests
 - Complementarity: Equivalence Principle, Redshift drift, ...
 - Synergies with other facilities, including ALMA, Euclid & SKA