OBSERVATIONAL **CONSTRAINTS ON** Varying-alpha Domain WALLS

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Spatial Variations of α

HIRES/Keck: $\frac{\Delta \alpha}{\alpha} = (0,57 \pm 0,11) \times 10^{-5}$ UVS/VLT: Much smaller...

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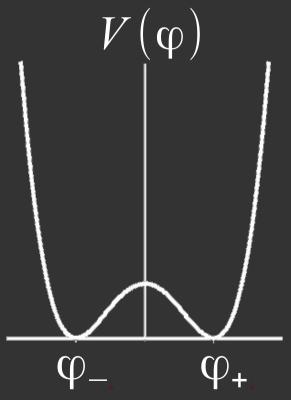
SPATIAL VARIATIONS OF lpha

COMBINED SAMPLE OF 295 ABSORBERS FROM VLT AND KECK IS COMPATIBLE WITH A DIPOLE MODEL WITH AMPLITUDE

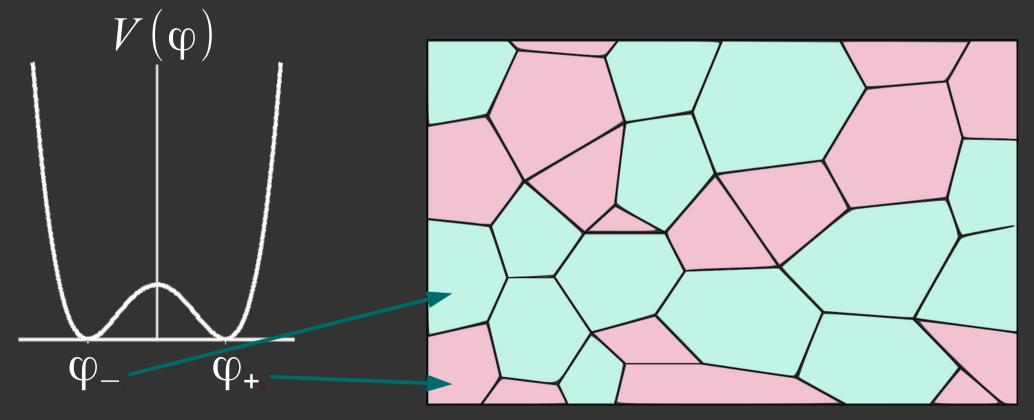
$$\frac{\Delta \alpha}{\alpha} = igl(0,97^{+0,22}_{-0,20} igr) imes 10^{-5}$$

[King et al, astro-ph:1202.4758]

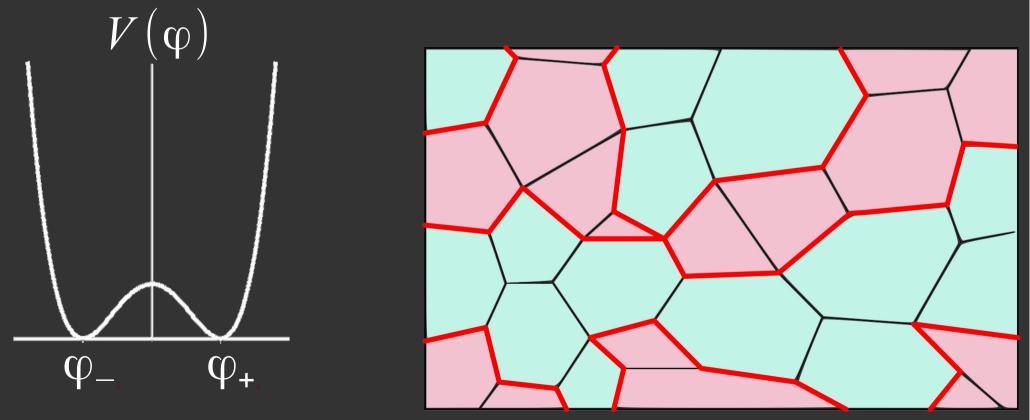
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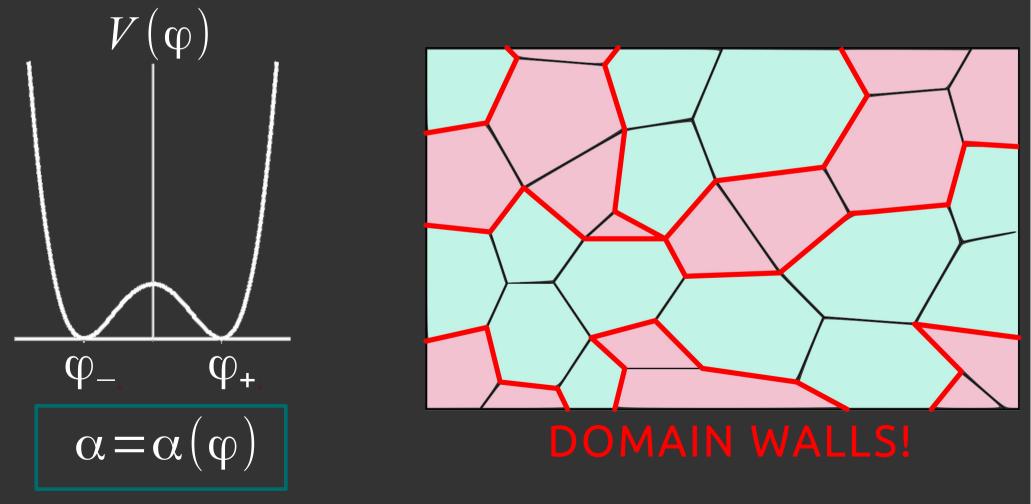


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DOMAIN WALLS!

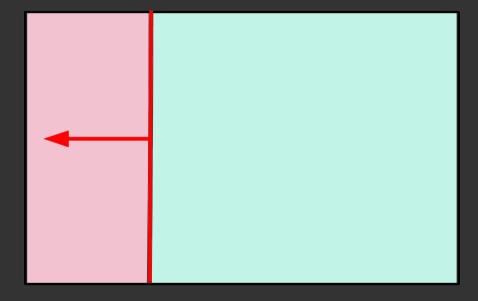
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DOMAINS ON EACH SIDE OF A WALL HAVE DIFFERENT BARYON ENERGY DENSITIES $\epsilon = \rho_{B+1} - \rho_{B-1}$

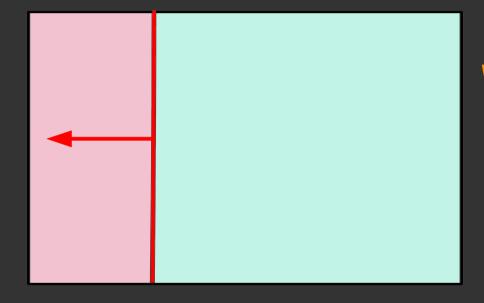


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DOMAIN WALLS ARE PUSHED TOWARDS REGIONS WITH A HIGHER ENERGY DENSITY.

IN COSMOLOGICAL BACKGROUNDS

$$\frac{dv}{dt} = \left|1 - v^2\right| \left|\frac{2}{R} + \frac{1}{\gamma R_{\epsilon}} - 3Hv\right|$$

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DECAY IS TRIGGERED WHEN BIAS TERM DOMINATES:

$$\frac{\sigma}{\epsilon L} = \frac{\Omega_w}{\Omega_B} \left| \xi \frac{\Delta \alpha}{\alpha} \right|^{-1} < 1$$

Observational Constraints

WALLS SHOULD BE AT $z \sim 0.5-1$ IN ORDER TO CAUSE REPORTED DIPOLE VARIATION OF α : [Olive et al, astro-ph:1204.4931]

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NARROW OBSERVATIONAL WINDOW

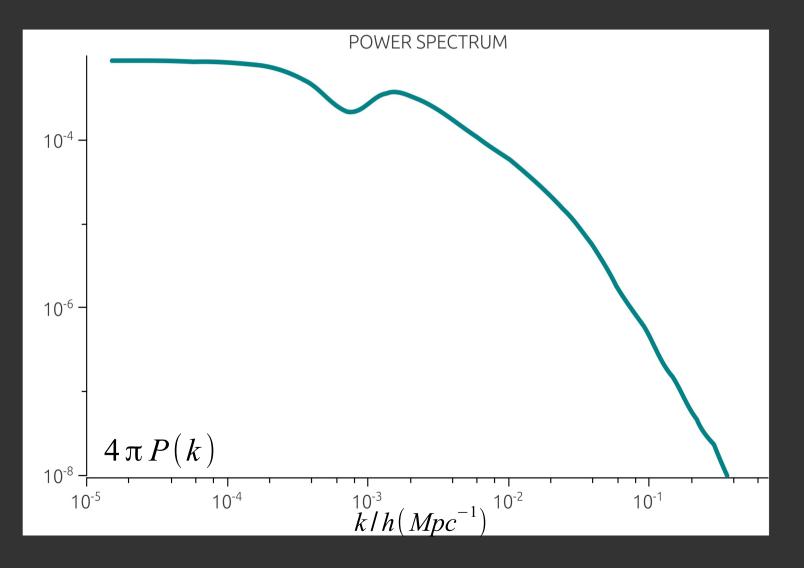
$$10^{-10} < \Omega_{w0} < 10^{-5}$$

CMB ANISOTROPIES

Spatial variation on α

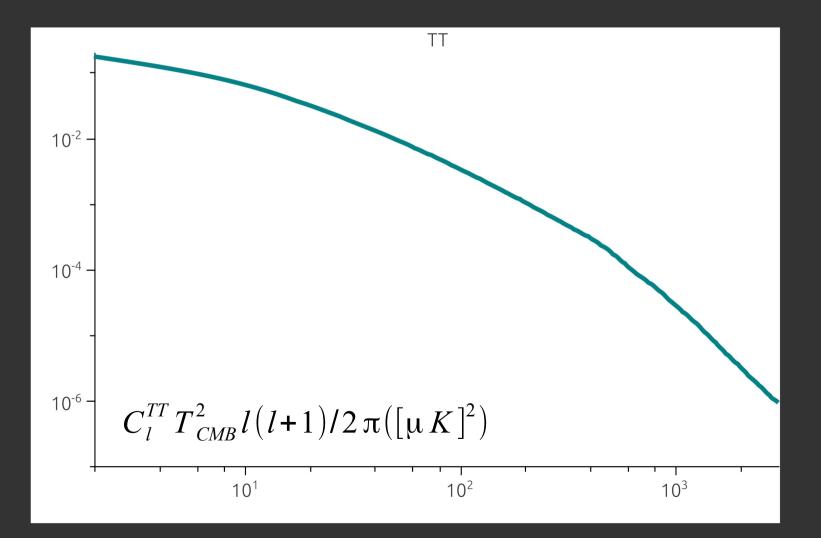
Coming Soon

CONTRIBUITIONS TO THE CDM POWER SPECTRUM ON SMALL SCALES IS REDUCED



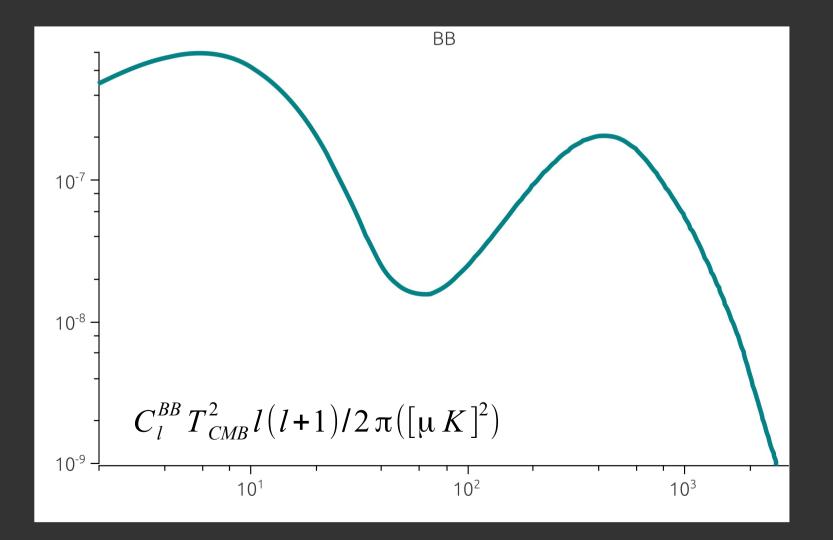
Coming Soon

UNLIKE OTHER DEFECTS CONTRIBUITIONS TO PRIMARY CMB ANISOTROPIES ARE NEGLIGIBLE



Coming Soon

HOWEVER, CONTRIBUITIONS TO B-MODE POLARIZATION MAY STILL BE SIGNIFICANT.



AKNOWLEDGMENTS

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P. P. Avelino and L. Sousa, *Observational Constraints of Varying-alpha domain walls*, Universe, 1, 6-16, 2015

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