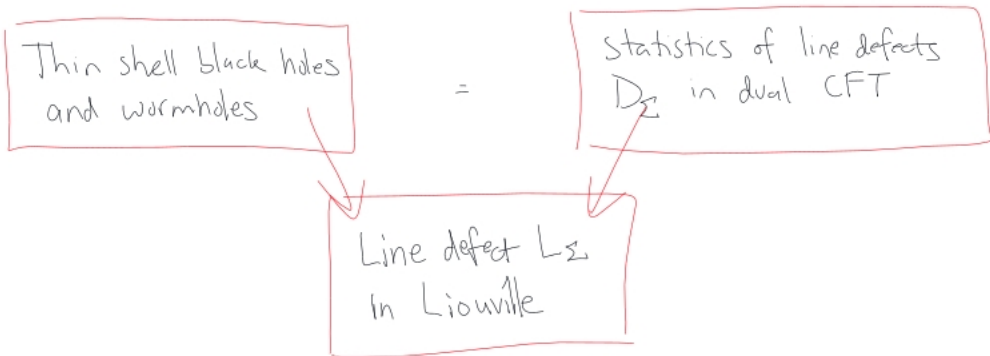


Statistics of 3D Black Holes from Liouville Line Defects

2206.03414 w/ Chandra
2404.15183 w/ Chandra, Merliya

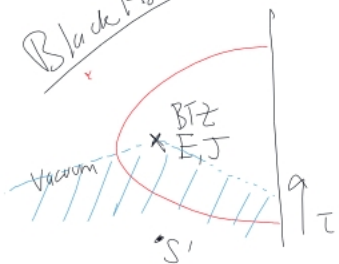


Gravity

$$- \int \sqrt{g} (R + 2) + m \int dy dl$$



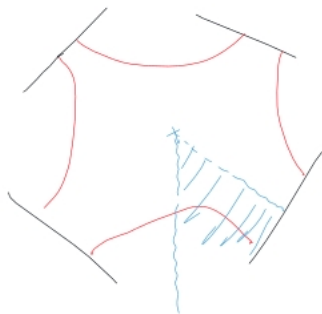
Black Hole



$$= \langle D_{\Sigma}^+ D_{\Sigma,1} \rangle$$

$$D_{\Sigma} |0\rangle = |4\rangle$$

Wormholes



$$= \frac{\langle D_{\Sigma}^+ D_{\Sigma} \rangle^4}{\leftarrow ??}$$

Q: How to reproduce this in dual CFT?

"Identity Block" \rightsquigarrow Cardy/Tauberian theorem

\rightsquigarrow reinterpret as ensemble

Liouville

$$S_L = \int \partial\phi\bar{\partial}\phi + e^\phi$$

$$V_\alpha = e^{2\alpha\phi}$$

$$L_\Sigma = e^{\frac{m}{4\pi} \int_\Sigma \phi}$$

Goal #1: find $\langle \mathcal{P} | L_{\Sigma} | \mathcal{P}' \rangle$

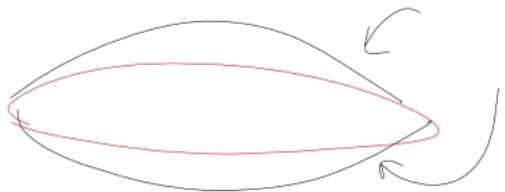
$C \rightarrow \infty$: classical solutions

$$\partial \bar{\partial} \phi = \frac{1}{2} e^{\phi}$$

$$\Delta[\partial_n \phi] = -2m$$

$$e^{\phi} |dz|^2$$

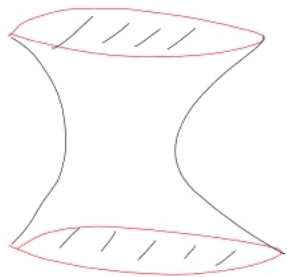
$$\langle L_{\Sigma} \rangle_{\text{sphere}} =$$



$$e^{\phi} = \frac{1}{\sinh^2(y \pm A)}$$

$$\langle L_{\Sigma}^{(+)} L_{\Sigma}^{(-)} \rangle_{\text{sphere}} =$$

$$\downarrow \langle 0 | L_{\Sigma} | P \rangle$$



$$\approx e^{-\frac{c}{6} S_{\text{on-shell}}}$$

$$\langle L_{\Sigma}^{+} L_{\Sigma} \rangle_{\text{torus}}$$

$$\hookrightarrow \langle P' | L_{\Sigma} | P \rangle = e^{\frac{c}{6} \boxed{C_L(h, h')}}_{\text{known}}$$

On disk:



boundary condition

$$e^{\phi} \sim \frac{4}{-(z-\bar{z})^2}$$

$$\langle 0 | L_{\epsilon} | z\bar{z} \rangle$$

$$= \int dP \underbrace{\langle 0 | L_{\epsilon} | P \rangle}_{\text{known}} \underbrace{\langle P | z\bar{z} \rangle}_{\text{known}}$$

Dual CFT₂

any $c > 1$

* line defect D_Σ that transforms like L_Σ (m)

$$* D_\Sigma^+ D_\Sigma \supset \mathbb{1}$$

Claim: universal matrix elements @ high E

$$|\langle i | D_\Sigma^{(m)} | j \rangle|^2 \stackrel{\text{States}}{\approx} C_0^D(h_i, h_j; m) C_0^D(\bar{h}_i, \bar{h}_j; m)$$

Ex @ large c

$$\lim_{n \rightarrow \infty} \prod_{i=1}^n \theta(z_i)$$

$$e^{\frac{m}{4\pi c} \phi} = \lim_{n \rightarrow \infty} \prod_{i=1}^n V_{i,n}$$

Derivation of $C_0^D(0, h; m)$



$$F_{\uparrow}^D = N \langle 0 | L_{\Sigma} | z z \rangle \quad (\text{exact})$$

universal matrix elements

$$C_0^D(h, h') \sim \langle h' | L_{\Sigma} | h \rangle e^{-\frac{1}{2}(S(h) + S(h'))}$$

↙ Cardy

$$L_{\Sigma}^+ L_{\Sigma} \sim |z z \rangle \langle z z|$$

$$D_{\Sigma}^+ D_{\Sigma}$$

Back to Gravity

$D_\Sigma \leftrightarrow$ dust shell

claim: dual to an approx. Gaussian

ensemble of $\langle i | D_\Sigma | j \rangle$

with variance $|C_0^D|^2$

Black holes

Vacuum



$$dp^2 + \cosh^2 p e^{\phi} |dz|^2$$

solution



$$\begin{aligned} \langle D_{\Sigma}^{\dagger} D_{\Sigma} \rangle &= Z_{\text{grav}} = e^{-2S_L} \\ &= |\langle 0 | L_{\Sigma} | z z \rangle|^2 \\ &= \langle D_{\Sigma}^{\dagger} D_{\Sigma} \rangle | \downarrow \downarrow \end{aligned}$$

Wormholes



Then

expand

$$|\psi\rangle = D_{\Sigma}^+ |0\rangle = \sum_P d_p |p\rangle$$

$$Z_{\text{wormhole}} = e^{-I_{\text{grav}}} = \overline{\langle \psi | \psi \rangle \langle \psi | \psi \rangle \langle \psi | \psi \rangle}$$

$$\text{with } \overline{d_p d_q} = \delta_{pq} C_0^D(0, h_p) C_0^D(0, \bar{h}_p)$$

Comments

* higher D

* In UV-complete theories?

$$Z(\tau, \bar{\tau}) Z(\tau, \bar{\tau})$$

Higher topologies \longrightarrow destabilized?

\searrow canceled?

UV (no wormholes)

{ time
average?
↓

{ state
average?
↓

{ RG
(Maldacena-Gi)
(Almheiri-Lin)
↓

{ large-N
jitter
↓

EFT (yes wormholes)